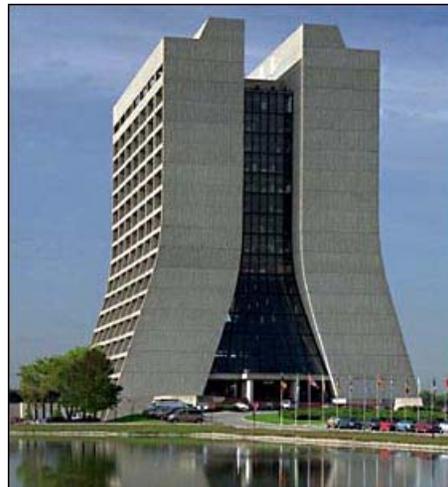




# Computational Cosmology in Chicago

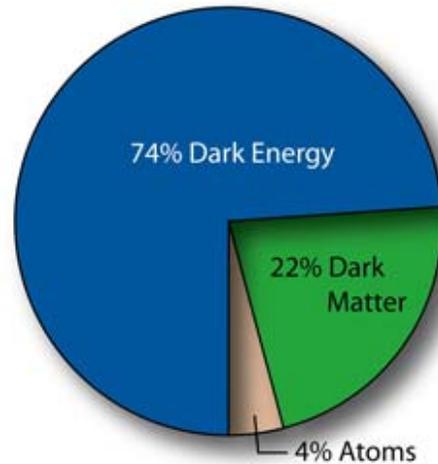
**Nick Gnedin**

Theoretical Astrophysics Group  
Fermilab





# Dark Energy



*Dark energy ranks as one of the most important discoveries in cosmology, with profound implications for astronomy, high-energy theory, general relativity, and string theory.*

Dark Energy Task Force Report



# Dark Energy Landscape

- **Stage II:**  
SNLS, ESSENCE, **SDSS-II**, CfASP, NSF, KAIT, CSP, QUEST, HST, PanSTARRS-1, PISCO, **SPT**, ACT, XCS, RCS2, DLS, KIDS, DEEP2
- **Stage III:**  
**DES**, HETDEX, WFMOS, PanSTARRS-4, ODI, 1000PLS, ALPACA, CIX, CCAT
- **Stage IV:**  
LSST, JDEM (**SNAP**, DESTINY, JEDI, ADEPT), SKA



# Dark Energy Probes

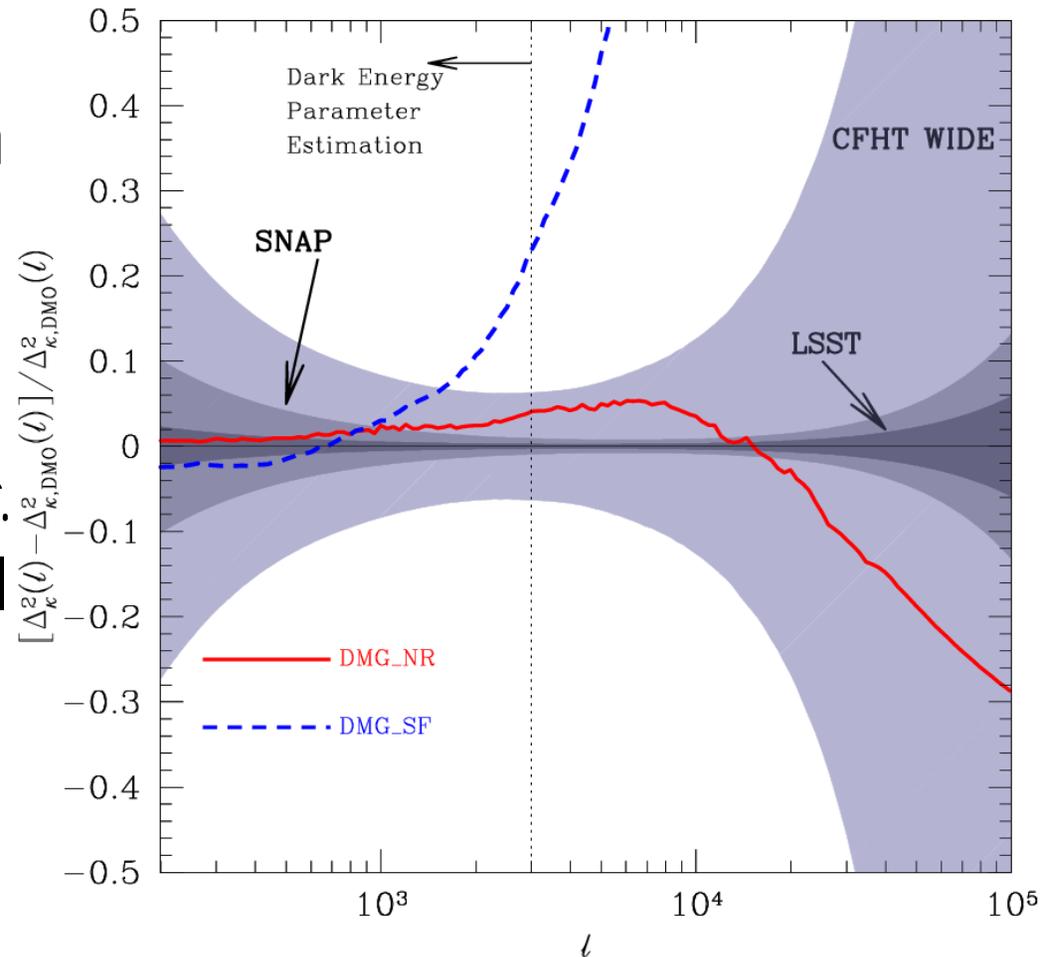
- **Supernovae Ia:**  
standard candles, luminosity distance vs redshift
- **Weak Lensing:**  
several measures of evolving matter distribution
- **Baryon Acoustic Oscillations:**  
standard ruler, angular-diameter distance vs redshift
- **Clusters of Galaxies:**  
number counts vs redshift



# Computational Cosmology for Fundamental Physics

## Weak lensing + BAO:

- 😊 Directly probe the evolution of the dark energy.
- 😞 Unfortunately, we exist.
- 😊 *Baryonic effects can be calibrated by simulations.*
- 😞 The required computational effort is massive.



(Rudd et al 2007)



# Computational Cosmology for Astrophysics

## Galaxy Formation:

- 😊 Direct comparison with most observations.
- 😞 Important physics (such as star formation) will not be understood any time soon.
- 😊 There exist simple scaling laws in the ISM.
- 😞 Models based on those laws make wrong galaxies.





# Computational Cosmology in the World



- **Europe:** Research mostly done by large groups:
  - Virgo Consortium [ $>30$  people] (Germany + UK)
  - Project Horizon [ $>20$  people] (France)
- **USA/Canada:** Large number of small groups (10-12 people)  
Chicago (FNAL+KICP+ANL), UWashington, Harvard, CITA, SLAC, LANL, Princeton, UCLA, UIUC, Berkeley, UMass...



# Computing Needs

- Modern state-of-the-art cosmological simulations require in excess of 1,000,000 CPU-hours (100 CPU-years). Simulations of 3,000,000 CPU-hours are not far away (and will be needed for calibrating DE experiments).
- A single simulation produces a Terabyte-scale data set.
- Even rudimentary analysis requires intermediate-level computing capability.

*The time of "this is a workstation to analyze your simulation" is over.*



# Solution (a-la LQCD)

- National consortium/collaboration (the level of integration may vary)
- Powerful local, intermediate-level resources (10,000 CPU, spread over several places: FNAL, SLAC, ANL, LANL, ...)
  - Analysis and visualization
  - Launching pad for national centers (30,000 CPU scale; development, porting, scaling)
- United approach to funding agencies and national supercomputer centers.



# Our Plan

- build-up local resources on par with other US groups
- create a Chicago-wide collaboration (FNAL, KICP, ANL)
- convince higher-ups
- overtake other US groups in local computing resources (reach 1,000 – 2,000 cores; there are good reasons why FNAL is the best place for that)
- move towards a national consortium/collaboration:
  - join with DOE labs (LANL, LBL, SLAC)
  - invite universities to join
- get lots of money along the way... (people, software support)



# 5-month Update

- FRA Collaborative Award (FNAL+UC)
- ANL LDRD awarded for 3 years
- 560 cores by the end of the year at FNAL
- $\frac{1}{4}$  ANL FTE working on data managements
- $\frac{1}{4}$  ANL FTE working on visualization
- Learning from FLASH
- A joint ANL/FNAL postdoctoral fellow to work on DES simulations (Fall 2008)
- Lots of proposal writing...